Corrections.

1. In a paper by the undersigned, "The TiO Colour Effect, and the Densities of M Stars" (T. P. $28._5$, 1936) the concluded cosmic spread in the density logarithm of the M dwarfs is based on an erroneously computed probable error of the parallax logarithm. Instead of $\Delta_{log\,\pi} = \pm 0.063$, there should be $\Delta_{log\,\pi} = \pm 0.0274$.

With this value of $\Delta_{log\,\pi}$, the cosmic spread in the density logarithm becomes $\Delta = \pm 0.15$, or ± 41 per cent.

This value of Δ, although amounting to more than twice the originally published value, is not large enough to make necessary considerable changes in the qualitative conclusions arrived at in the above-mentioned paper.

2. A similar error occurs in the paper: "On the Empirical Mass-Luminosity Relation" (T. P. $30._1$, 2, 1938). There the cosmic spread of the masses from the mean mass-luminosity relation was considered. Instead of $\Delta_{log\,\pi} = \pm 0.04$, there should be $\Delta_{log\,\pi} = \pm 0.017$. However, this correction has no influence upon the final conclusion, since the dispersion now becomes equal to the observational error dispersion, instead of amounting to only two-fifths of it, and the cosmic spread in the mass logarithms for given luminosity remains practically zero, as was already suggested in the original paper.

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Correction.

In a paper by E. Öpik, "Stellar Structure etc" (T. P. $30._3$, 1938), p. 51, Table 2, $n = 0.0$, for $Q_{max}/A = 0.309$ read 0.103, and for $\alpha = 0.057$ read 0.171; the error is produced by the omission of a factor of $\frac{1}{3}$. As the case $n = 0.0$ is nowhere used in the above-mentioned paper, the error is of no consequence.